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The SP19 chronology for the South Pole Ice Core - Part 2: gas chronology, Δage, and smoothing of atmospheric records

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Abstract. A new ice core drilled at the South Pole provides a 54,000-year paleoenvironmental record including the composition of the past atmosphere. This paper describes the SP19 chronology for the South Pole atmospheric gas record and complements a previous paper (Winski et al., 2019) describing the SP19 ice chronology. The gas chronology is based on a discrete methane (CH_4) record with 20- to 190-year resolution. To construct the gas time scale abrupt changes in atmospheric CH_4 during the glacial period and centennial CH_4 variability during the Holocene were used to synchronize the South Pole gas record with analogous data from the West Antarctic Ice Sheet Divide ice core. Stratigraphic matching based on visual optimization was verified using an automated matching algorithm. The South Pole ice core recovers all expected changes in CH_4 based on previous records. Smoothing of the atmospheric record due to gas transport in the firn is evident but relatively minor, despite the deep lock-in depth in the modern South Pole firn column. The new gas chronology, in combination with the existing ice age scale from Winski et al. (2019), allows a model-independent reconstruction of the gas age-ice age difference through the whole record, which will be useful for testing firn densification models.

1 Introduction

Ice core records provide detailed reconstructions of past climate in the polar regions and unique global records of the past atmosphere. They are valuable recorders of past climate because very accurate time scales can be created for both the gas and ice phase, allowing precise comparisons to events in other ice cores and paleoarchives (e.g., Elderfield et al., 2012; Hodell et al., 2017; Marcott et al., 2013). The recently collected South Pole ice core (SPC14) expands a spatial array of ice cores drilled across Antarctica that extend into the last glacial period.

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SPC14 is an intermediate depth (1751 m) ice core that was drilled as a part of the South Pole ice core (SPICEcore) project and was collected in the 2014/15 and 2015/16 field seasons (Souney et. al., 2020). The core provides ice and gas data through part of last glacial period, to 54,302 years before present (BP, with 0 BP = 1950 CE; Winski et al., 2019). Because drilling stopped almost 1000 m above bedrock, folding and mixing of layers at the bottom of the core is not a concern, resulting in a stratigraphically continuous record for the entire length of the core. The core location is 89.99˚ S, 98.16˚ W, at surface elevation of 2835 m on the polar plateau of the East Antarctic ice sheet. The current annual accumulation rate is 8 cm/a ,water equivalent, (Lilien et al., 2018; Mosley-Thompson et al., 1999) with an annual-mean temperature of -51˚ C as measured in the firn (Severinghaus et al., 2001). Due to its geographic location, ice accumulating at the site has low levels of trace impurities (Casey et al., 2017). These characteristics are an advantage for the measurements of ultra-trace gases (for example, ethane, methyl chloride, and methyl bromide; Aydin et al., 2004; Saltzman et al., 2004; Nicewonger et al., 2018) one of the primary goals of the SPICEcore project.

Air permeates through firn and is trapped at depth between 50-120 m, depending on local accumulation and temperature. As a result, ice and trapped air at the same depth are different in age, with the air being younger (Schwander and Stauffer, 1984). The gas age-ice age difference (Δage) depends on the ice accumulation rate and temperature, and can range from tens to thousands of years. Gas chronologies for previously collected ice cores have been created either through calculating gas ages using an existing ice chronology coupled with models of Δage, or by stratigraphically matching features in gas records with previously dated records in other cores (Schwander and Stauffer, 1984; Sowers et al., 1992; Schwander et al., 1997; Petit et al., 1999; Blunier et al., 2007; Bazin et al., 2012; Veres et al., 2013; Buizert et al., 2014). In cold, relatively low accumulation-rate sites similar to the South Pole, Δage model uncertainty can be a major contributor to the overall uncertainty in the gas age time scale. At the South Pole site today, Δage at the bubble close-off depth is ~1000 years, large enough that the classical approach to calculating Δage using a firn-densification model, typically having a ~20% uncertainty, is insufficient for dating the gas at the precision needed to compare of leads and lags between abrupt climate signals recorded in the ice core.

This paper focuses on the creation of a CH_4-based gas chronology for SPC14 using a stratigraphic matching approach and is a companion to a paper describing the first ice chronology for the core. The gas and ice chronologies are collectively referred to as the SP19 chronology. CH_4 is a well-mixed atmospheric trace gas exhibiting globally synchronous abrupt variations on decadal to millennial scales (Blunier and Brook, 2001; Brook et al., 1996; Lee et al., 2018; Rhodes et al., 2015), making it an ideal choice for stratigraphic matching to existing ice core records. The chronology presented here relies on correlating CH_4 variations between the SPC14 and the West Antarctic Ice Sheet Divide (WD) ice cores, using millennial-scale abrupt variations during the last ice age and glacial-interglacial transition, and centennial-scale variations during the Holocene. We first describe relevant attributes of the SPC14 core and acquisition of the SPC14 CH_4 record, and then we discuss synchronization and optimization of the gas chronology. We also discuss key observations from our results, including implications for the gas age-ice age difference (Δage), smoothing of atmospheric gas records by gas transport in the South Pole firn, and short-term variability in atmospheric CH_4.

2 Methods

2.1 CH_4 measurements

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High-resolution CH_4 concentration measurements were made along the entire length of SPC14, jointly at Oregon State University (OSU) and Pennsylvania State University (PSU). Samples from the 139 - 1077 m interval were measured at PSU and samples in the 1078 - 1751 m interval were measured at OSU. Both labs measured samples in the 330 – 840 m and 1130 - 1150 m intervals for intercalibration. PSU concentrations were increased by 6 ppb to correct for an offset that was revealed by the intercalibration measurements. A total of 2318 measurements (733 at PSU and 1598 at OSU) were made on samples at 1067 individual depths resulting in 1 to 2-meter depth resolution throughout the entire core. Samples were measured in duplicate (832 depths), triplicate (46 depths), or in quadruplicate for the purpose of laboratory intercalibration (109 depths). 80 sample depths were measured without replication due to limited sample size or poor sample quality.

CH_4 concentrations measured at OSU were made using a wet-extraction technique as described in Grachev et al. (2009), with updates by Mitchell et al. (2011) and Lee et al. (2018). Briefly, subsamples of the main core measuring 10 cm x 6 cm x 2.5 cm (with the 10 cm dimension oriented parallel to the vertical axis of the core) were split into replicate samples by cutting along the vertical axis. Each individual sample was then placed in a glass vacuum flask and attached to an automated analytical setup. The samples were kept frozen by immersing the flasks in an ethanol bath at -68 ˚C. After evacuating atmospheric air from the flasks with a vacuum pump, the flasks were submerged in a warm water bath for 30 minutes, melting the ice samples and releasing the trapped air. The water was then refrozen, equilibrating to the temperature of the ethanol bath, over a period of 1 hour. Once the temperature of the flasks stabilized to -68 ˚C, air in the head space of each flask was expanded four times into a gas chromatograph (GC) for CH_4 analysis. Concentrations were quantified by comparison to a calibrated air standard at the beginning and end of each day (500.22 ppb for samples measured in 2016 and 481.25 ppb for any samples measured after 2016) on the NOAA04 CH_4 concentration scale (Dlugokencky et al., 2005).

Several corrections were made to the raw CH_4 concentration value measured at OSU including adjustments for a small quantity of CH_4 that remains dissolved in the melt water (Mitchell et al., 2013; Lee et al., 2019); because gases do not reach complete solubility equilibrium during the melt-refreeze process, an empirical solubility correction is employed. Mitchell et al., (2013) describes the experimental derivation of the correction. The derivation was repeated for SPC14 samples resulting in a correction factor of 1.7 %. All sample concentrations measured at OSU were corrected for solubility by increasing the measured value by 1.7%.

A small amount of CH_4 can also be present in measured samples due to the influence of air leaks or other contamination. An additional blank correction was applied to OSU measurements to account for this small contamination. To quantify the blank correction, air-free ice (AFI) was routinely measured in conjunction with samples. Production of AFI is described by Mitchell et. al (2013). AFI was processed for analysis and measured in the same way as sample ice; however, an amount of standard air with a known mole fraction of CH_4 was added to the flask with AFI prior to the melt-refreeze step. Blank corrections derived from these measurements, also corrected for solubility effects, were subtracted from the measured concentration of each sample. Because samples were measured at different times, a blank correction was applied to each group of samples depending on when they were measured. Average blank corrections ranged from 6.6 ppb to 9.8 ppb for all samples measured at OSU. Data and information about corrections are provided in the supplementary material.

PSU CH_4 measurements (depths 139 m – 1077 m) were also made using an automated melt-refreeze method similar to the OSU system. However, the PSU system uses stainless steel flasks, which introduces an additional blank correction associated

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with CH_4 outgassing. This blank correction was estimated to be 35 ppb, and was applied to all PSU samples. Further description of the PSU method can be found in WAIS Divide Project Members, (2013).

All CH_4 concentrations are slightly affected by fractionation in the firn column due to gravity (Craig et al., 1988; Mitchell et al., 2011; Schwander et al., 1997; Sowers et al., 1992). The amount of gravitational fractionation is controlled by the thickness of the diffusive zone in the firn column, which can be estimated by measuring the $\delta^{15}N$ of N_2. We corrected all measured CH_4 concentrations for gravitational fractionation with $\delta^{15}N$ of N_2 data (Winski et al., 2019) by interpolating the $\delta^{15}N$ of N_2 to the depths of the CH_4 samples, and then using the relationship (Mitchell et al., 2011):

$$CH_{4corr} = CH_{4\,meas} \times \left(1 + \Delta M \frac{\partial^{15}N}{1000}\right) \quad (1)$$

where ΔM is 12.92 g/mol, the difference between the mass of air (M = 28.96 g/mol) and the mass of the CH_4 (M = 16.04 g/mol). $\delta^{15}N$ varies from 0.63‰ to 0.46‰, with a mean value of 0.54‰. The correction ranges from 2.7 to 5.7 ppb (1 σ = 0.83 ppb).

The SPC14 discrete CH_4 record, measured jointly at OSU and PSU, spans 130 years to 52,482 years BP. Sample spacing of the CH_4 measurements is between 20 - 190 years, increasing with depth. CH_4 concentrations vary from 355 ppb to 751 ppb. Pooled standard deviation for the measurements is 2.9 ppb between 130 m to 1150 m is 2.9 ppb, and 2.7 ppb between 1150 m – 1751 m. The record resolves CH_4 signals observed in previous ice cores (Fig.s 1 and 7). The mean difference between the reference CH_4 record from WAIS Divide (WD) and the SPC14 CH_4 records, determined by interpolating WD CH_4 data to the ages of SPC14 CH_4, samples is only 2.9 ppb ± 0.96 (one standard deviation; n = 1067), demonstrating the long-term stability of the measurement systems.

2.2 Gas Chronology

2.2.1 Summary of Synchronization Approach

To create a gas chronology for SPC14, CH_4 variations were visually matched at equivalent rapid CH_4 variations in the WD ice core; subsequently the match was optimized using an automated algorithm. The rapid changes during the last glacial period are coincident with the Northern Hemisphere Dansgaard-Oeschger events (Baumgartner et al., 2014; Huber et al., 2006; Rosen et al., 2014; Severinghaus and Brook, 1999; Severinghaus et al., 1998) and are excellent chronostratigraphic tie points between the ice cores. The SPC14 CH_4 record also resolves the abrupt CH_4 features associated with Heinrich events, as described by Rhodes et al. (2015) and further resolves centennial scale variations in CH_4 previously described in the WD (Mitchell et al., 2013) and Roosevelt Island (RICE) ice cores (Lee et al., 2019), and in several records by Rhodes et al. (2017). The centennial variations are smaller in magnitude than the D-O events but are clearly present and are used as Holocene tie points in the SPC19 gas chronology (Table 1, Fig. 7).

Synchronization of rapid CH_4 excursions between ice core records requires that both records are adequately sampled. WD was chosen as the basis for the SPC14 gas time scale because of (1) its accurate and precise chronology (WD2014) based on annual layer counting and CH_4 ties to Greenland ice cores and speleothem chronologies (Buizert et al., 2015; Sigl et al., 2016); (2) its high resolution continuous (Rhodes et al., 2015) and discrete (Mitchell et al. 2013; WAIS Divide Members, 2015) CH_4 record, minimally smoothed by gas transport in the firn (Buizert et al., 2015; Sigl et al., 2016); and (3) volcanic matching between the SP14and WD cores, providing a South Pole ice chronology synchronized to WD2014 (Winski 2019). We used the WD discrete CH_4 record for 0 - 9.8 ka BP, and the continuous CH_4 record for 9.8 ka BP until 54 ka BP (Rhodes et al., 2015; WAIS Divide Project Members, 2015). The WD2014 chronology has also been used with success for synchronization of other Antarctic ice cores (Buizert et al., 2018; Lee et al., 2018). The resulting SPC14 CH_4 record has an age resolution of 25 to 150 years, which is sufficient for resolving all of the major abrupt CH_4 variations of the last 54,000 years as well as smaller-scale Holocene variations.

2.2.2 Tie Point Selection and Gas Age Uncertainty

Matching CH_4 variations between WD and SPC14 records establishes the WD2014 gas age at the depth of the SP14CH_4 feature being matched. Because the SP19ice chronology has been volcanically synchronized to WD2014, this also allows us to empirically establish Δage at the depth of the CH_4 feature. The full gas chronology is then constructed by interpolating Δage between these tie points using a cubic spline, and subtracting the spline from the ice chronology.

Tie point selection was done in two stages, first by visual matching, followed by fine-tuning of the visual match using an automated optimization algorithm. We first visually selected either the midpoint, maximum, or minimum of abrupt changes in CH_4, depending on the shape of the event, as tie-points. The midpoints of D-O and Heinrich CH_4 events were determined by averaging CH_4 before and after each abrupt change, then determining the midpoint between these averages, using the same techniques for averaging and defining the midpoint as described in Buizert et al., (2015). We then optimized the tie points using a best-fit algorithm that randomly perturbs the age of each visually selected point within a 500-year window centered around the visual tie point. The tie points were perturbed individually (i.e., one at a time). Each tie point age was randomly perturbed 1,000 times, after which the goodness-of-fit was calculated by finding the minimum misfit, using equation (2):

$$S_m = \frac{1}{2}\sum(g_m - g_o)^2 \tag{2}$$

where S_m is the misfit, g_m are the SPC14 CH_4 data after perturbing a tie point, and g_o are the data we match to (i.e. the methane record of WD on the WD2014 chronology); we apply a linear interpolation to find the g_0 at the exact same ages as the g_m. Once the best tie point for that event was found, the iteration was performed on the next older tie point. The automated optimization was done on high-pass filtered versions of the CH_4 records (1st order Butterworth with 50-year cut off), thereby eliminating any bias created by low frequency measurement offsets.

The procedure resulted in a final tie point selection where the adjustment ranged from 0.3 years to 189 years (with a mean change of 23 years) from the hand selected tie point, giving confidence that the matching is robust. Correlation of the WD and SPC14 high-pass filtered records increased from r = 0.9599 (visual matching) to r = 0.9634 (automated matching). The final tie points are listed in Table 1. In the supplemental material gas ages are listed for all depths provided in the SPC14 ice age time scale data file (Winski et al., 2019) to provide unified SP19 ice age and gas age time scales for future use. Both time scales are plotted in Fig. 2.

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Three factors impact the uncertainty of the resulting gas chronology. The first is correlation uncertainty, i.e., how accurately the age of the tie point is transferred from WD to SPC14. This uncertainty is primarily controlled by the sample spacing around each tie point. The second factor is uncertainty that arises from the cubic-spline interpolation between tie points, which is more difficult to quantify. The cubic spline interpolation used here eliminates discontinuities at tie points but is not representative of the physical processes of firn densification and layer thinning. To estimate interpolation uncertainty, we examined the agreement of small-scale methane variations between the tie points that were not explicitly matched in the procedure. Based on this evaluation the interpolation uncertainty is up to 106 years in the Holocene and up to 190 years in the glacial period. A continuous estimate of the interpolation uncertainty requires that we account for the increase in the uncertainty with distance from a tie point. Based on Fudge et al. (2014) we allow the interpolation uncertainty to increase to 10% from the distance of the closest tie point. The third factor to consider is the absolute uncertainty of the reference (WD) gas chronology (Buizert et al., 2014), which incorporates uncertainties in the WAIS Divide ice age time scale and WAIS Divide Δage model. To find the estimated 2 σ uncertainty along the length of the core, we used the root sum square of all three uncertainties. The uncertainties are provided in the supplement and shown in Fig. 1a.

3 Results and Discussion

3.1 An empirical record of Δage for SPC14

Accurately constraining the gas age-ice age difference (Δage) is critical for interpreting ice core records. Traditionally, Δage is calculated using firn densification models (Arnaud et al., 2000; Barnola et al., 1991; Goujon et al., 2003; Loulergue et al., 2007; Lundin et al., 2017; Schwander et al., 1997). These models simulate the physical process of firn densification over time to determine the depth and age (relative to the surface) of trapped air. Input parameters for the models (temperature, accumulation rate, surface snow density, close-off density) as well as the physical processes involved in densification, are not known well in many cases, leading to substantial uncertainties in Δage when estimated through a model. This is particularly a problem in locations or past time periods where Δage is relatively large. The difficulty in simulating past firn densification has led to difficulties in interpretation of the relative phasing of greenhouse forcing and Antarctic climate (Brook and Buizert, 2018).

SPC14 has independent ice and gas chronologies, synchronized to the WD2014 chronology via volcanic and CH_4 markers, respectively. The independently dated ice and gas chronologies allow us to compute an empirically derived Δage history for SP14with a very low relative uncertainty due to the fact that WD has a small Δage (and therefore also a small absolute Δage uncertainty). The SPC14 ice chronology was created by combining annual-layer counting with stratigraphic matching of volcanic events, and is annually resolved through the Holocene (Winski et al., 2019). Uncertainty for the Δage record is impacted by three factors, including: (1) the WD Δage uncertainty, (2) correlation uncertainty between chosen CH_4 tie points in the record, and (3) uncertainty in the ice age interpolation between volcanic tie points. These terms were added in quadrature to estimate a 2 σ uncertainty for the empirical SPC14 Δage record, which increases with age (Fig. 3).

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The Δage record (Fig. 3) is the first of its kind for Antarctica. It shows the expected larger Δage during the glacial period than the Holocene (due to both lower temperatures and lower accumulation rates) and an overall increase from 55 to 25 ka associated with the cooling from Marine Isotope Stage (MIS) 3 to MIS2. To assess the origin of the Holocene Δage variations, we compare our new empirical Δage to a preliminary firn densification model results presented in Winski et al. (2019). This preliminary model is forced with temperature derived from $\delta^{18}O_{ice}$ and accumulation rates based on the annual-layer identification corrected for ice flow thinning (Fig. 4) which are only available only for the Holocene section. The modelled and empirical Δage show the same pattern of variations and agree within 5% back to 6 ka and within 10% back to 8 ka, suggesting that the Holocene Δage variations we reconstruct are realistic and likely driven by changes in accumulation rate at the site. The data-model comparison of Fig. 4 suggests that the Δage record presented here is a good target for benchmarking firn densification models, which are important for describing the physical firn densification process.

3.2 Smoothing of the SPC14 atmospheric gas record

Due to the slow firn densification process, gas diffusion and gradual bubble formation in the firn column act as a low-pass smoothing filter on the atmospheric signal (Buizert et al., 2013; Fourteau et al., 2017; Gregory et al., 2014; Schwander et al., 1993). As the firn densifies, pores remain largely open to the atmosphere, allowing the atmospheric gases to diffuse freely. At the lock in depth (LID), the firn begins to close off and diffusion of air stops (Battle et al., 1996a; Kawamura et al., 2006; Mitchell et al., 2015). Once pore close-off occurs, no more mixing with the air above can occur. Although the impact of smoothing in the firn on gas records has long been recognized, it is not well quantified because it depends on physical processes near the firn-ice transition that are difficult and time consuming to study (Fourteau et al., 2019).

The degree to which the atmospheric signal as recorded in the ice has been filtered is of interest for understanding the speed of past environmental changes, the fidelity of the ice core gas record, and also impacts gas-to-gas correlation like the technique employed here. For example, in a situation where an abrupt CH_4 increase were heavily smoothed, the damping of the concentration change (Spahni et al., 2003) would impact a tie point location. At the South Pole this issue could be a concern because this site has an unusually deep lock in depth (LID), currently ~110 m (Battle et al., 1996; Severinghaus and Battle, 2006).

To quantify the preservation of the SPC14 CH_4 signal at specific abrupt events we compared prominent CH_4 features between the SP14and WD cores. A comparison of event duration in the WD core and the difference in amplitude between the event in WD and SPC14 is presented in Table 2 and Fig. 5. Event duration was determined by the number of years between the onset of rapid increases in CH_4 and when CH_4 returned to pre-event levels. As expected, our results indicate that the amplitude of shorter lived events is reduced more than longer lived events (amplitude reduction varies from 0 to 31 ppb), consistent with the findings of Spahni et al. (2003) who examined the smoothing of the 8.2 ka methane event in the EPICA Dome C ice core. However, even at values of Δage approaching 2400 years (Fig. 3), which are reached during the last glacial period, previously identified fast CH_4 variations are still preserved faithfully in SP14(Fig. 5). This level of preservation gives us confidence not only in the accuracy of the tie points, but also in how well other atmospheric gas records will be preserved in SPC14.

We apply a simple model approach to further examine how much smoothing has affected the SPC14 record. We start with the WD methane record as input, apply various smoothing filters (gas age distributions) based on a firn model, and compare the

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results to the SPC14 record. In doing so we assume that the WD record is a reasonable substitute for the true atmospheric history; this assumption is justified by the high accumulation rate in WD and narrow age distribution (Mitchell et al., 2015).

Site smoothing is fully described by the gas age distribution in the closed bubbles. The gas age distribution employed here was created using a firn air transport model tuned to modern-day Dome C firn air sampling data and site conditions that incorporates advection, diffusion, near-surface convective mixing, deep firn dispersion and gradual bubble trapping (Buizert et al., 2012; Buizert and Severinghaus, 2016). The model was calibrated to the EDC site because it is the closest modern-day analogue to South Pole glacial conditions, with accumulation rates of around 3 cm a^{-1} and a Δage of around 2300 years. The resulting age distribution is similar to that obtained by Spahni et al. (2003) using a similar approach.

The spectral width Δ of the gas age distribution is defined as (Trudinger et al., 2002):

$$\Delta^2 = \frac{1}{2}\int_0^{\infty}(t-\Gamma)^2 G(t)dt$$

with G the gas age distribution in yr^{-1} and Γ the mean of the distribution. The spectral width of the simulated EDC present-day closed-bubble gas age distribution equals 78 years (corresponding to around 3.5% of EDC Δage today).

In our analysis we assume that the spectral width of the gas age distribution scales linearly with Δage, or $\Delta = \alpha\times\Delta$age; where α is unitless scaling factor. This is a reasonable assumption one can make about the system, given that Δage represents the timescale of the snow-to-ice transformation; the gradual bubble trapping that dominates the broadening of the age distribution likely scales with this process to a large degree.

We seek to quantify smoothing in the SPC14 CH_4 record by estimating the optimal scaling parameter α. We filter the WD CH_4 record (assumed to reflect the true atmospheric variations) by a gas age distribution that is a linearly scaled version of the simulated EDC distribution – scaled such that its spectral width reflects $\alpha\times\Delta$age at that given time in the core. We repeat this exercise for a wide range of α values from 1×10^{-2} to 1 in 100 equally spaced steps. The newly filtered WD CH_4 record is then compared to the SP CH_4 record to determine the optimal value of α that best represents the observed degree of smoothing by minimizing a misfit function. This is illustrated in Fig. 6 (where α is expressed as a percentage rather than a fraction).

The best fit to the SPC14 record uses a smoothing function history with a spectral width of 3% of Δage (or $\alpha = 0.03$) The data and our analysis show clearly that despite the large values of Δage, significant short-term variability will be preserved at ice core sites like the South Pole.

3.3 Centennial variations in CH_4

The SPC14 record validates previous observations of persistent centennial-scale CH_4 variability through the Holocene and in the glacial period (Mitchell et al., 2013; Rhodes et al., 2017; Lee et al., 2018) including variations matched with the WD CH_4 record back to 16,150 ka, just after the onset of the glacial termination (Fig. 7). The centennial scale features observed during the Holocene are important for understanding pre-anthropogenic CH_4 variations. Atmospheric CH_4 variations in the last 2,000 years have been attributed to anthropogenic forcing mechanisms (Ferretti, 2005; Mischler et al., 2009; Sapart et al., 2012). However, recent work on the Roosevelt Island ice core (RICE) and WD ice cores and now the new SPC14 record (Fig. 7) validate the existence of similar CH_4 variations beginning as early as the last glacial period, well before the influence of anthropogenic forcing (Lee et al., 2018; Rhodes et al., 2017). The observation of centennial scale variations throughout the Holocene implies that these small but consistent CH_4 variations occur naturally, rather than caused exclusively by human activity (Lee et al., 2018), though their origin remains unclear. Rhodes et al., (2017) hypothesized that they represent small changes in low-latitude climate conditions, which lead to small changes in methane production, although whether they are forced or arise as a consequence of internal variability is an open question.

4.0 Summary and conclusions

The SP19 gas chronology for the SPC14 ice core is presented for the last 52,482 years, complementing the ice chronology presented in Winski et al. (2019). The gas chronology was created using over 2,000 high resolution, discrete CH_4 measurements completed at Oregon State University and Pennsylvania State University. The resulting CH_4 record was tied to the high resolution CH_4 record of the WAIS Divide ice core using the WD14 chronology. Abrupt changes in CH_4 at D-O events as well as distinct variations of 20-30 ppb during the Holocene are used as tie points. The absolute uncertainty of the gas chronology changes through time to a maximum (1 σ) of ± 540 years at 35 ka, and an uncertainty of ± 502 years at the bottom of the core. Key outcomes of this study include a gas age time scale for the SPC14 ice core, the observation of minimal smoothing of the gas record despite the exceptionally deep firn column at the South Pole, an empirical Δage record that can be used to test firn densification models, and the confirmation of centennial variability in atmospheric CH_4.

Data Availability

The data are available in the supplementary material and the post-review time scale and data will be made fully available at the NOAA National Center for Environmental Information Paleoclimate Data base (https://www.ncdc.noaa.gov/data-access/paleoclimatology-data) and the USAP Antarctic Glaciological Data Center (http://www.usap-data.org).

Author Contributions

All authors contributed data to this study. JE, EB, CB, JSE, TS, JS, EH ad MK measured ice core gases. EK and ES made isotope measurements. DW, EO, TF, KK, DF, and JK measured ice core chemistry and contributed to the ice chronology which was used to calculate delta age. JE, EB, and CB created gas chronology. DW, TJF, DF, EK, MA oversaw the ice core collection. JE, EB, and CB wrote the manuscript with input from all authors.

Competing Interests

The authors declare that they have no conflict of interest.

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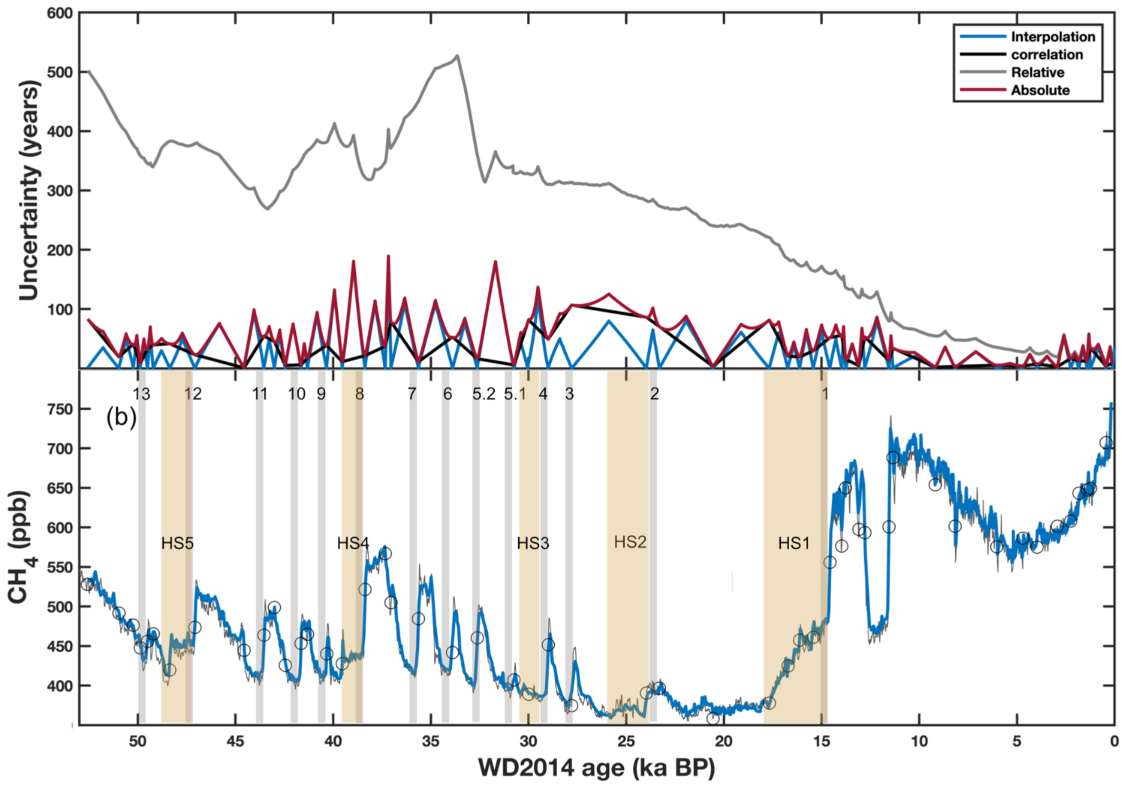

Figure 1. a) SP19 gas chronology uncertainty. Black line indicates correlation uncertainty, blue line is interpolation uncertainty, red solid line is total SPC14 uncertainty relative to WD. The red line is a combination in quadrature of the correlation and interpolation uncertainty. The grey line describes total absolute uncertainty, which incorporates the absolute uncertainty in the WD time scale. Maximum uncertainty of ± 540 years is found around 35 ka. b) SPC19 methane record (blue line), plotted on top of WD CH_4 record (grey) (Rhodes et al., 2015; WAIS Divide Project Members, 2015). Selected tie points are indicated by circles. The gas chronology extends from 116 years to 52,482 years BP. Grey shaded bars indicate D-O events, yellow bars indicate Heinrich Stadials.

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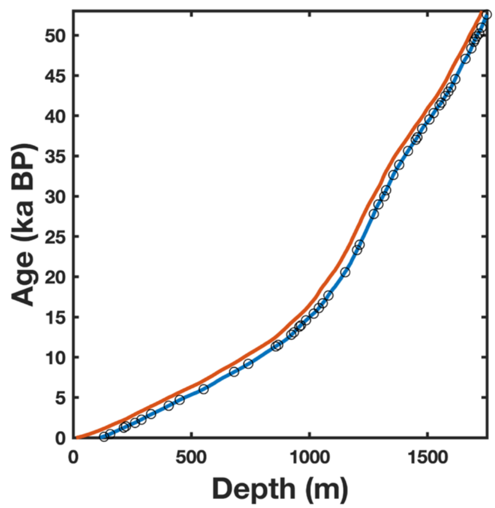

Figure 2. Ice age (orange) and gas age (blue) as a function of depth for the SP19 chronology. Gas tie points are indicated by black circles.

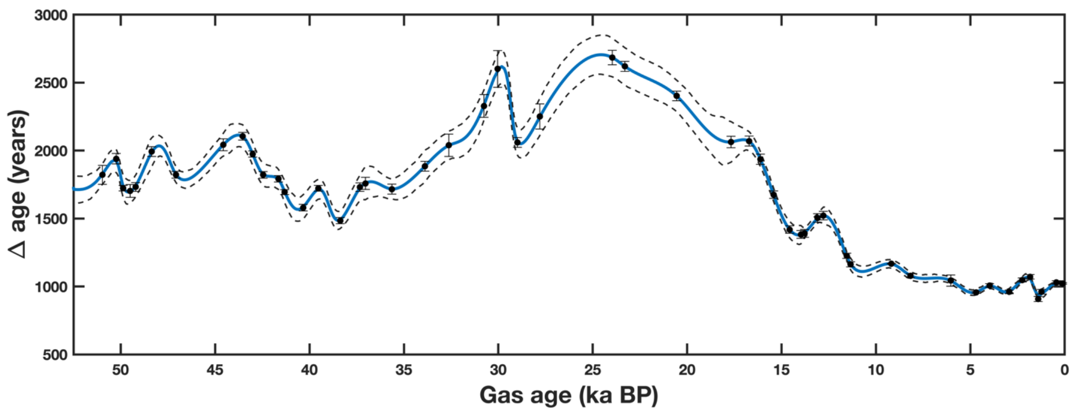

Figure 3. Empirically derived Δage history for the SPC14 ice core. Blue line is a spline fit to the Δage points. Δage error bounds (2 σ), dashed lines, reflect uncertainties with Δage based on WD Δage uncertainty, and relative SP19 uncertainties, black dots indicate individual Δage constraints (see text).

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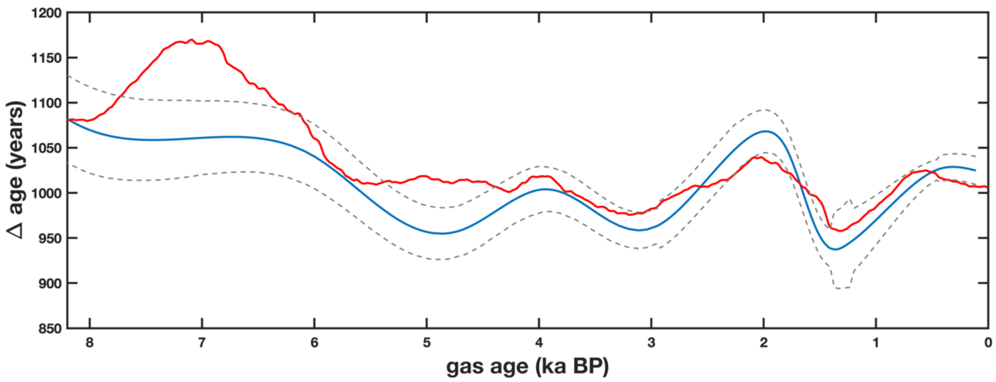

Figure 4. Comparison of modelled Δage (red; see text for details) and empirical Δage (blue line). Grey dashed lines represent Δage uncertainty.

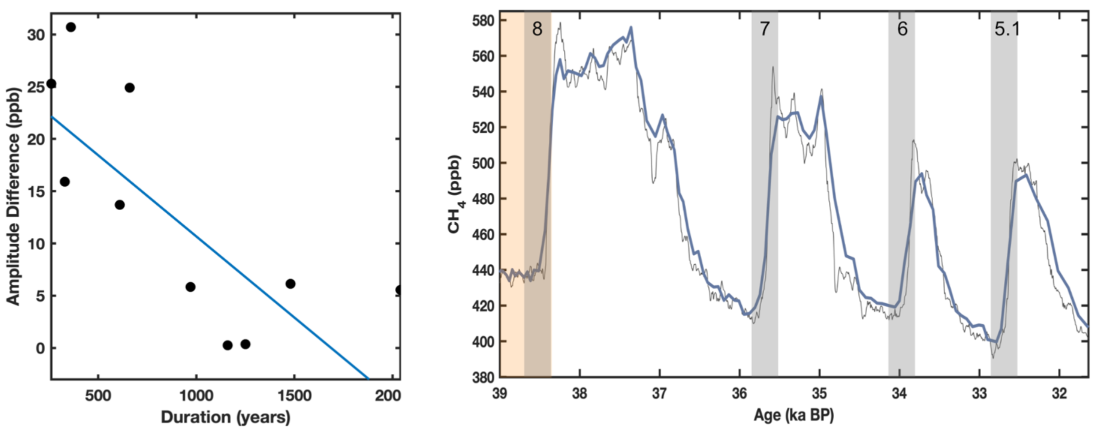

Figure 5. (left) Correlation of duration of event and amplitude difference between WD and SPC14 events shows a clear negative trend (r = -0.76, p = 0.0066). (right) Example of smoothing from MIS3 showing smoothing of small-scale features in SPC14 (blue) relative to WD (grey). Grey bars indicate D-O event, Heinrich Stadial 4 is shaded in orange.

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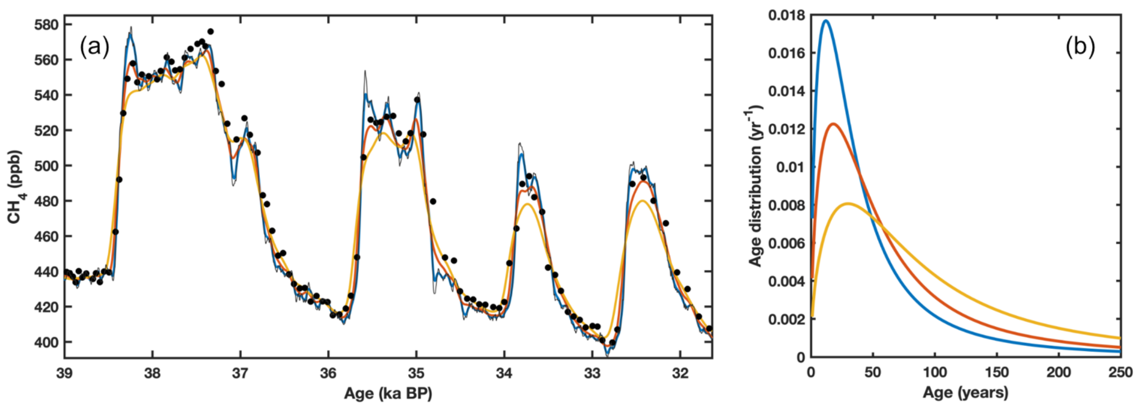

Figure 6. a) SPC14 data (black dots) compared with smoothed WD CH_4 record. Colored lines show the result of smoothing the WD record with iteratively wider age distributions. The width of the age distributions corresponds to the median age of the distribution. Original signal (grey) is plotted against three example smoothed histories: median age of 1 % (blue), median age of 3.0% (red), and median age of 5% (orange) of Δage. The best fit between the smoothed WD record and SPC14 data is 3.0 % of Δage. b) Width of the smoothing filter is defined by the median age, which is proportional to a percentage of Δage. Colors of the filter correspond with smoothed signal in (a).

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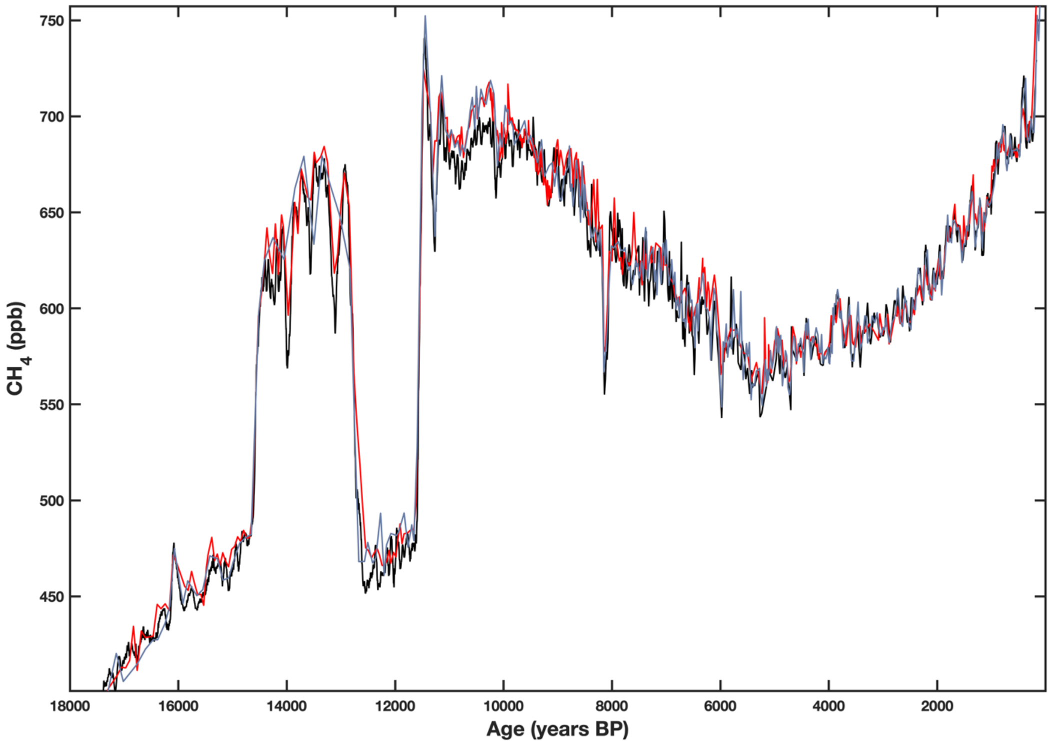

Figure 7. a) SPC14 (blue), WD (black), and RICE (red) all exhibit centennial variation in CH_4 in the Holocene and during the deglaciation. RICE data from (Lee et al., 2018), WD data from (Rhodes et al., 2015).

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Table 1: Tie points used for chronology with Δage and uncertainty. Uncertainties are listed at the tie points as correlation uncertainty and interpolation uncertainty. Interpolation uncertainty is given as the largest estimate for an interval between tie points. Δage is reported in years at each tie point. See supplementary material for complete time scale uncertainties. 1950 CE = 0 years.

SPC14 Depth (m)	Gas age (yr)	Δage (yr)	Correlation uncertainty (± yr)	SPC14 Depth (m)	Gas age (yr)	Δage (yr)	Correlation uncertainty (± yr)
130.20	113	955	5.71	1379.45	33889	1899	35.95
156.25	443	1021	12.78	1417.25	35635	1714	36.64
214.20	1233	960	17.39	1450.00	37036	1744	44.60
223.60	1406	919	18.76	1457.02	37334	1725	31.36
261.22	1831	1052	15.95	1476.78	38368	1472	22.24
288.50	2250	1040	15.19	1507.95	39537	1728	20.82
330.00	2938	993	13.90	1525.23	40337	1584	23.75
403.91	3978	1020	13.58	1550.00	41314	1679	17.50
450.56	4690	956	5.75	1558.81	41640	1790	23.56
552.00	6039	1041	41.62	1575.65	42454	1801	23.69
681.06	8168	1079	14.54	1588.15	43007	1982	23.12
741.33	9183	1190	1.71	1599.17	43541	2101	28.32
857.20	11341	1154	18.69	1616.50	44564	2049	43.77
867.77	11547	1224	19.25	1660.05	47077	1835	25.02
923.32	12783	1541	31.63	1684.42	48367	2008	35.88
934.36	13092	1496	31.13	1695.79	49206	1733	32.47
956.36	13778	1404	29.25	1700.66	49503	1694	47.37
963.13	13969	1372	29.16	1707.16	49873	1716	19.92
985.21	14570	1409	26.97	1718.27	50231	1751	38.16
1017.50	15421	1657	27.68	1727.50	50969	1860	70.03
1039.86	16121	1935	37.69	1751.00	52586	1768	44.70
1056.00	16713	2073	36.60				
1080.51	17677	2020	42.65				
1151.00	20558	2365	35.12				
1200.79	23300	2639	37.14				
1212.87	23963	2662	53.03				
1271.87	27798	2295	92.17				
1291.23	28992	2039	36.09				
1316.30	30039	2351	134.89				
1324.07	30756	2356	83.98				
1354.49	32624	2032	82.31				

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Table 2: Illustration of smoothing of the methane record in SPC14. Table shows comparison of event duration to the amplitude difference of events in the SPC14 and WD core (see Fig. 5). Age resolution of SPC14 samples for the duration of the event are also given.

Event Name	Amplitude difference (ppb)	Event Duration (years)	SPC14 resolution (years)
YD-onset	0.1	1250	37
DO-3	16	330	55
DO-4	31	360	40
DO-5	25	660	66
DO-6	14	610	47
DO-7	6	1480	78
DO-8	6	2040	93
DO-9	25	260	52
DO-10	5	970	139
DO-11	0.2	1160	58
DO-12	-3	1360	68

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